



Introduction to the Covariant Notation:

All of the ideas and equations presented in this course are in the Lorentz invariant (relativistic invariant) notation usually referred also as covariant notation. This is for two reasons:

- The first reason addresses the need of relativistic invariant theories. One of the most fundamental principles in physics is the fact that all equations and subsequently their predictions should be invariant with respect to the frame of reference. Otherwise physics would not be an objective science.
- The second reason is a practical one but no less important and it addresses the need for using special relativity in High Energy Physics (HEP): At the energies that the HEP deals most particles have velocities approaching the velocity of light so they need to be treated according to the formulae of the Special Relativity. So we start by introducing the covariant notation and the Lorentz transformations within this notation:

Every point in space-time can be represented by a *contravariant 4-vector* defined as:

$$\mathbf{x}^\mu = (\mathbf{x}^0; \vec{\mathbf{x}}) = (ct; \vec{\mathbf{x}}) \quad (1)$$

The vector index in (1) is running between $\mu = 0, 1, 2, 3$ and the 4-vector is defined explicitly as:

$$\mathbf{x}^\mu = (\mathbf{x}^0; \mathbf{x}^1, \mathbf{x}^2, \mathbf{x}^3) = (ct; \mathbf{x}, \mathbf{y}, \mathbf{z})$$

Greek indices denote always space-time variables whilst latin indices denote always 3-dimensional variables. A more rigorous definition of a 4-vector in terms of its transformation properties will be given later once the Lorentz transformation has been defined in covariant notation.

Define to dot-product of two contravariant vectors to be the matrix product:

$$\mathbf{x} \cdot \mathbf{y} = \mathbf{x}^\mu \mathbf{g}_{\mu\nu} \mathbf{y}^\nu \quad (2)$$

Where :

$$\mathbf{g}_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad (3)$$



is the metric tensor of a flat Minkowski space. Several books including Perkins use several different metric definitions. The end result does not of course change but it does create confusion. The definition in (3) is the most popular in HEP and Relativity books and this is what we will follow throughout this course.

It is worth noticing that:

(1) We have used the Einstein convention where two indices that are the same indicate summation. This of course implies that in a given expression one cannot have more than 2 indices which are the same. Sometimes we refer to this summation as index contraction because the summed indices disappear at the end. This is also why summed indices can be re-named at wish: Since they disappear you can change the index to whatever symbol is convenient for the calculation.

(2) The metric indices are subscripts in this case whilst the contravariant vector indices are always superscripts. This is not an accident will become clear later why we do that.

In a more explicit way the definitions (2) (3) mean that:

$$\mathbf{x} \cdot \mathbf{y} = x^\mu g_{\mu\nu} y^\nu = \begin{pmatrix} x^0 & x^1 & x^2 & x^3 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} y^0 \\ y^1 \\ y^2 \\ y^3 \end{pmatrix} = x^0 y^0 - \vec{\mathbf{x}} \cdot \vec{\mathbf{y}}$$

Using the definitions the dot-product of a 4-vector with itself is:

$$(1)(2)(3) \Rightarrow \mathbf{x} \cdot \mathbf{x} = (x^0)^2 - \vec{\mathbf{x}}^2 = (ct)^2 - \vec{\mathbf{x}}^2.$$

Hence **the dot-product of a 4-vector with itself is a relativistic invariant** since a spherical light wave should look spherical in every coordinate frame. It will be shown later that the dot-product of *any* 4-vectors is relativistic invariant.

Although we can do all of calculations using contravariant vectors we will need to include always the metric when we need to take dot-products. This is inconvenient and this is why we need to introduce the *covariant vectors*:

A covariant vector is defined as:

$$x_\mu = (x^0; -\vec{\mathbf{x}}) = (ct; -\vec{\mathbf{x}}). \quad (4)$$

Notice that except the space vector sign which is negative we have also changed the 4-vector index from superscript to subscript.



Covariant tensors are always associated with subscripts in this notation. The dot-product can now be defined from the covariant and contravariant vectors without the explicit use of the metric as:

$$\mathbf{x} \cdot \mathbf{x} = \mathbf{x}_\mu \mathbf{x}^\mu = (ct)^2 - \vec{\mathbf{x}}^2 = \mathbf{g}_{\mu\nu} \mathbf{x}^\nu \mathbf{x}^\mu \quad (5)$$

Equations (4) and (5) imply that:

$$\mathbf{x}_\mu = \mathbf{g}_{\mu\nu} \mathbf{x}^\nu$$

As seen here **the metric can be used to lower an index and convert a contravariant vector to a covariant vector**. The opposite is also true if one defines the metric to be the same for both covariant and contravariant indices:

$$\mathbf{g}^{\mu\nu} = \mathbf{g}_{\mu\nu}$$

and in this case the metric can be used to rise an index:

$$\mathbf{x}^\mu = \mathbf{g}^{\mu\nu} \mathbf{x}_\nu$$

and convert a covariant 4-vector to a contravariant 4-vector.

In this notation one can define the Kroneker delta as:

$$\delta^\mu_\nu = \mathbf{g}^{\mu\rho} \mathbf{g}_{\rho\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Since it must be that

$$\delta^\mu_\nu = \frac{\partial \mathbf{x}^\mu}{\partial \mathbf{x}^\nu}$$

we are driven to the convention where contravariant indices in the denominator become covariant indices in the numerator and visa versa. This can be seen also in the case where

$$\mathbf{g}^{\mu\nu} = \frac{\partial \mathbf{x}^\mu}{\partial \mathbf{x}_\nu}$$



Clearly in this notation we have that $\mathbf{g}_{\mu\nu}\mathbf{g}^{\nu\mu} = 4$.

Contravariant and covariant derivatives are then defined as:

$$\partial_{\mu} = \frac{\partial}{\partial \mathbf{x}^{\mu}} = \left(\frac{\partial}{\partial \mathbf{x}^0}; \nabla \right)$$

and

$$\partial^{\mu} = \frac{\partial}{\partial \mathbf{x}_{\mu}} = \left(\frac{\partial}{\partial \mathbf{x}^0}; -\nabla \right)$$

Lorentz Transformations

Our definition of a contravariant 4-vector in (1) whilst easy to understand is not the whole story. **A contravariant 4-vector is an object defined as $\mathbf{x}^{\mu} = (\mathbf{x}^0; \vec{\mathbf{x}})$ that transforms as a vector under Lorentz transformations.** That is:

$$\mathbf{x}^{\mu'} = \Lambda^{\mu}_{\nu} \mathbf{x}^{\nu} \quad (6)$$

where \mathbf{x}^{ν} is the 4-vector in a frame \mathbf{O} and $\mathbf{x}^{\mu'}$ a 4-vector in the reference frame \mathbf{O}' related by the Lorentz transformation Λ^{μ}_{ν} . The matrix Λ^{μ}_{ν} is in the general case a complex object which can represent a mixture of Lorentz boosts and 3D rotations. However, in the case of a Lorentz boost by $\vec{\beta} = \beta \hat{\mathbf{x}}_0 = (\mathbf{v}/c) \hat{\mathbf{x}}_0$ in the x-direction it assumes a form that is familiar from special relativity texts which can be written as:

$$\Lambda^{\mu}_{\nu} = \begin{pmatrix} \gamma & -\beta\gamma & \mathbf{0} & \mathbf{0} \\ -\beta\gamma & \gamma & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{1} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{0} & \mathbf{1} \end{pmatrix} \quad (7)$$

where $\vec{\mathbf{v}}$ is the relative velocity of the frame \mathbf{O}' with respect to \mathbf{O} .



$$(6)(7) \Rightarrow \mathbf{x}^{\mu'} = \begin{pmatrix} x^{0'} \\ x^{1'} \\ x^{2'} \\ x^{3'} \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix} \Rightarrow$$

$$\begin{aligned} x^{0'} &= \gamma(x^0 - \vec{\beta} \cdot \vec{x}) \\ x^{1'} &= \gamma(x^1 - \beta x^0) \\ x^{2'} &= x^2 \\ x^{3'} &= x^3 \end{aligned} \quad (8)$$

The four equations in (8) represent a Lorentz boost for $\vec{\beta} = \beta \hat{x}_0$.

Now that the general form of a Lorentz transformation has been defined under (6) we can investigate the consequences of the relativity constraint that the product of a 4-vector with itself must be Lorentz invariant:

$$\begin{aligned} \mathbf{x}' \cdot \mathbf{x}' &= \mathbf{g}_{\mu\nu} x^{\mu'} x^{\nu'} = \mathbf{g}_{\mu\nu} \Lambda^\mu_\alpha x^\alpha \Lambda^\nu_\beta x^\beta \Rightarrow \\ \mathbf{x}' \cdot \mathbf{x}' &= (\mathbf{g}_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta) x^\alpha x^\beta = \mathbf{x} \cdot \mathbf{x} \Rightarrow (\mathbf{g}_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta) = \mathbf{g}_{\alpha\beta} \Rightarrow \\ (\Lambda^\mu_\alpha \Lambda_{\mu\beta}) &= \mathbf{g}_{\alpha\beta} \Rightarrow \mathbf{g}_{\alpha\beta} \mathbf{g}^{\beta\sigma} = (\Lambda^\mu_\alpha \Lambda_{\mu\beta}) \mathbf{g}^{\beta\sigma} \Rightarrow \\ \Lambda^\mu_\alpha \Lambda_{\mu}^\sigma &= \delta_\alpha^\sigma \end{aligned} \quad (9)$$

This is a basic property of the Lorentz transformations. It is as simple exercise to show that (7) is consistent with (9).

Now we are ready to show that the product of any two 4 vectors is a Lorentz invariant quantity:

$$\begin{aligned} (8)(9) \Rightarrow \mathbf{x}' \cdot \mathbf{y}' &= x^{\mu'} y_{\mu'} = x^{\mu'} y^{\nu'} g_{\mu\nu} = \Lambda^\mu_\alpha x^\alpha \Lambda^\nu_\beta y^\beta g_{\mu\nu} \Rightarrow \\ \mathbf{x}' \cdot \mathbf{y}' &= \Lambda^\mu_\alpha \Lambda^\nu_\beta g_{\mu\nu} x^\alpha y^\beta = \Lambda^\mu_\alpha \Lambda_{\mu\beta} x^\alpha y^\beta = \Lambda^\mu_\alpha \Lambda_{\mu}^\beta x^\alpha y_\beta \Rightarrow \\ \mathbf{x}' \cdot \mathbf{y}' &= \delta_\alpha^\beta x^\alpha y_\beta = \mathbf{x} \cdot \mathbf{y} \end{aligned}$$



Hence, the **dot-product of any 4-vectors is a relativistic invariant**. Notice that this is always the product of a covariant with a contravariant 4-vector. So when we contract covariant with contravariant indices the results of the summation is always going to be a Lorentz invariant quantity which is what we usually want to have in Physics. It is for this reason that when we use Einstein's convention to sum over indices we always have two same indices but one is covariant and the other contravariant.

Next we can calculate the **inverse boost** of (6), $(\Lambda^{-1})^\alpha_\rho$. We start with (9) and multiply both sides by $(\Lambda^{-1})^\alpha_\rho$:

$$\Lambda^\mu_\alpha \Lambda^\beta_\mu = \delta_\alpha^\beta \Rightarrow \Lambda^\mu_\alpha (\Lambda^{-1})^\alpha_\rho \Lambda^\beta_\mu = \delta_\alpha^\beta (\Lambda^{-1})^\alpha_\rho = (\Lambda^{-1})^\beta_\rho \quad (10)$$

However since $(\Lambda^{-1})^\alpha_\beta$ is the inverse we have that:

$$\Lambda^\mu_\alpha (\Lambda^{-1})^\alpha_\rho = \delta^\mu_\rho \quad (11)$$

$$\delta^\mu_\rho \Lambda^\beta_\mu = (\Lambda^{-1})^\beta_\rho \Rightarrow \Lambda^\beta_\rho = (\Lambda^{-1})^\beta_\rho \quad (12)$$

The Lorentz transformation for vectors (6) can be extended for rank-2 tensors. **Hence, a rank-2 tensor is an object that under Lorentz transformations transforms as:**

$$\mathbf{T}^{\alpha\beta'} = \Lambda^\alpha_\mu \Lambda^{\beta'}_\nu \mathbf{T}^{\mu\nu} \quad (13)$$

It is then easy to show that the sum:

$$\mathbf{T}^{\alpha\beta} \mathbf{x}_\alpha \mathbf{x}_\beta \quad (14)$$

is a Lorentz invariant quantity if $\mathbf{T}^{\alpha\beta}$ transforms according to (13) and $\mathbf{x}_\alpha, \mathbf{x}_\beta$ transform according to (6). Obviously (13) can be extended to any rank-n tensor. In that case we have from (14) that any quantity where all the indices have been summed up (fully contracted indices) will be a relativistic/Lorentz invariant.



Parity Transformation Properties

A combination of the Lorentz and Parity properties can be used to classify the various fields and currents we use in Particle Physics as follows:

- Quantities which are invariant under Lorentz transformations are called *scalars* (**S**) if they are even under parity. If they are odd under parity they are called *pseudo-scalars* (**P**).
- Similarly by definition 4 vectors transform under parity as follows:

$$V^\mu = (V^0; \vec{V}) \rightarrow V_P^\mu = (V^0; -\vec{V})$$

and are commonly referred to as *vectors*.

- Another category of Lorentz Vectors transform differently under parity:

$$A^\mu = (A^0; \vec{A}) \rightarrow A_P^\mu = (-A^0; \vec{A})$$

and are called *axial-vectors*.



Maxwell's Equations in Covariant Notation

One can see how powerful can be the covariant notation by writing Maxwell's equations in a covariant form. Start with Maxwell's equations in the standard 3D notation:

$$\vec{\nabla} \cdot \vec{E} = 4\pi\rho$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

$$\vec{\nabla} \times \vec{E} = -\left(\frac{1}{c}\right) \frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J} + \left(\frac{1}{c}\right) \frac{\partial \vec{E}}{\partial t}$$

Define the electromagnetic 4-vector potential and 4 vector current to be

$$A^\mu = (\Phi; \vec{A}) \quad J^\mu = (\rho; \vec{J}) \quad (15)$$

where Φ, \vec{A} are the electromagnetic scalar and vector potential and ρ, \vec{J} are the charge and current densities. The Electromagnetic Field tensor can be defined as a function of the 4-vector potential as:

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu \quad (16)$$

which in terms of the electric and magnetic fields can be written as:

$$F^{\mu\nu} = \begin{pmatrix} \mathbf{0} & -E_x & -E_y & -E_z \\ E_x & \mathbf{0} & -B_z & B_y \\ E_y & B_z & \mathbf{0} & -B_x \\ E_z & -B_y & B_x & \mathbf{0} \end{pmatrix} \quad (17)$$

The field tensor can also be written as:

$$F_{\alpha\beta} = g_{\alpha\mu} g_{\beta\nu} F^{\mu\nu} = \begin{pmatrix} \mathbf{0} & E_x & E_y & E_z \\ -E_x & \mathbf{0} & -B_z & B_y \\ -E_y & B_z & \mathbf{0} & -B_x \\ -E_z & -B_y & B_x & \mathbf{0} \end{pmatrix} \quad (18)$$



Another form of the field tensor, $\tilde{F}^{\mu\nu}$, can be defined by:

$$\tilde{F}^{\mu\nu} = \frac{1}{2}\epsilon^{\mu\nu\alpha\beta}F_{\alpha\beta} \quad (19)$$

where $\epsilon^{\mu\nu\alpha\beta}$ is defined to be : $\epsilon^{\mu\nu\alpha\beta} =$ zero if any indices are the same, -1 for odd permutations and +1 for even permutations of the indices.

Using (15)(16)(17)(18)(19) the four Maxwell Equation can be written as two covariant equations:

$$\partial_{\mu}F^{\mu\nu} = \frac{4\pi}{c}J^{\nu} \quad (20)$$

$$\partial_{\mu}\tilde{F}^{\mu\nu} = 0 \quad (21)$$

The first one corresponds to the first and the fourth Maxwell equations which contain sources and the second one corresponds to the other two.